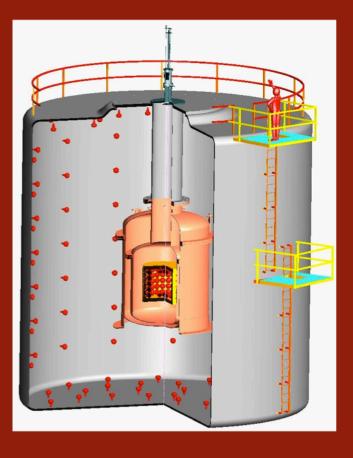
## The Coherent Elastic Neutrino Nucleus Scattering (CENNS) Experiment at the Booster Neutrino Beamline

PHYSICAL REVIEW D 89, 072004 (2014)

A method for measuring coherent elastic neutrino-nucleus scattering at a far off-axis high-energy neutrino beam target

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 (Received 25 November 2013; published 3 April 2014)



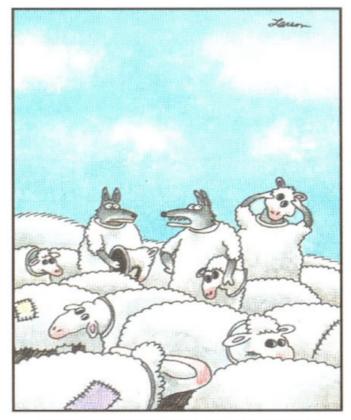
Robert Cooper

http://neutrino.indiana.edu/rlcooper



#### **Outline**

- Physics Motivation for CENNS
- How do we measure CENNS?
  - i.) Neutrino production
  - ii.) Detection
  - iii.) Background suppression
- The SciBath Detector
- Future work and conclusions



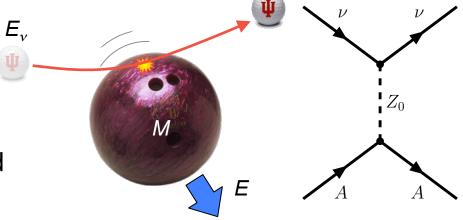
"Wait a minute! Isn't anyone here a real sheep?"

## **Describing the CENNS Signal**

To probe a "large" nucleus (few × 10<sup>-15</sup> m)

$$E_{\nu} \lesssim \frac{hc}{R_N} \cong 50 \text{ MeV}$$

Detector signature is the recoiling nucleus



Recoil energy that is deposited

$$E_r^{\rm max} \simeq \frac{2E_\nu^2}{M} \simeq 50 \text{ keV}$$

This is quite small for particle & nuclear physics → Dark Matter

## Structure of the CENNS Signal

Predicted scattering rate

$$\frac{d\sigma}{dE} = \frac{G_F^2}{4\pi} \left[ \frac{(1 - 4\sin^2\theta_w)Z - N}{2} M \left( 1 - \frac{ME}{2E_\nu^2} \right) F(Q^2)^2 \right]$$

 $\approx 0 \rightarrow$  protons have little influence

square of sum → part of coherence condition

nuclear form factor

→ distribution of neutrons

• Recoil energy  $(M^{-1})$  and rate  $(N^2)$ 

#### v Cross Sections vs. Energy Cross-section (10<sup>-38</sup> cm²) Coherent 10<sup>-2</sup> 10<sup>-3</sup> Bere be bragons 10<sup>-4</sup> 10<sup>-5</sup> electrons 10<sup>-6</sup> 70 80 90 100 Neutrino Energy (MeV) Image from K. Scholberg



## **Physics Cases for CENNS**

- Never been observed!
- Oscillations (spatially)
- Form factors
- Supernova physics
- Non-standard interactions
- Irreducible dark matter background



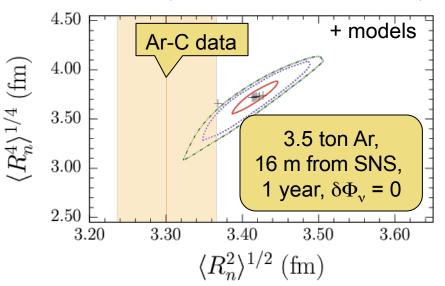
## **Physics Cases for CENNS**

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- Non-standard Interactions
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#### 4th vs 2nd Form Factor Moments

$$F(Q^2) = \frac{1}{Q_W} \left[ F_n(Q^2) - (1 - 4\sin^2\theta_W) F_p(Q^2) \right]$$

$$F_n(Q^2) \approx \int \rho_n(r) \left( 1 - \frac{Q^2}{3!} r^2 + \frac{Q^4}{5!} r^4 - \frac{Q^6}{7!} r^6 + \cdots \right) r^2 dr$$

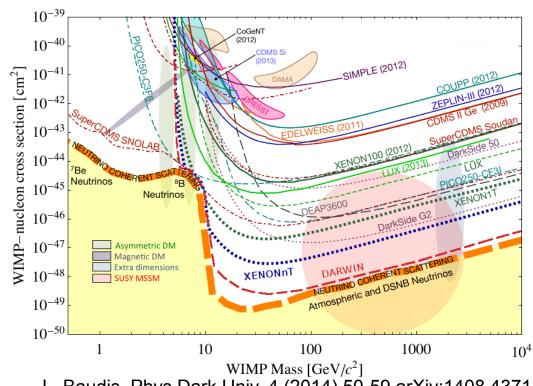


Patton et al., arXiv/1207.0693

## **Physics Cases for CENNS**

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#### **Dark Matter Sensitivity**

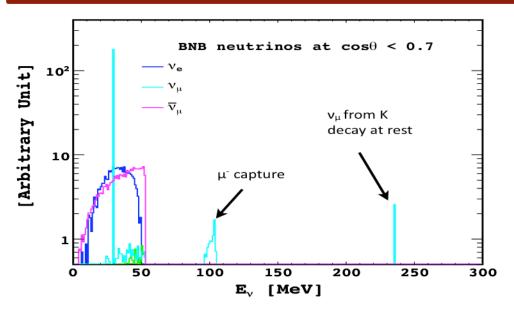


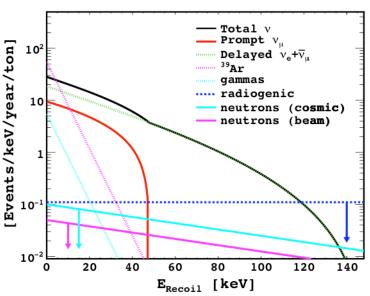
L. Baudis, Phys.Dark Univ. 4 (2014) 50-59 arXiv:1408.4371

#### MiniCLEAN for BNB CENNS

 A. Hime already gave an excellent review of MiniCLEAN for BNB CENNS measurement

½ ton LAr scintillation detector → 100 events / yr





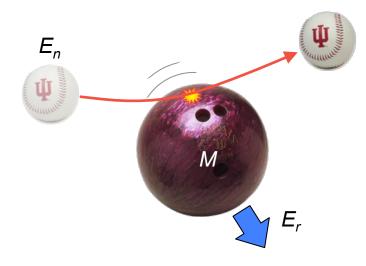
Inner

Vessel

## **Neutron Backgrounds**

- Few-MeV neutrons will deposit ~10 keV in LAr
- Accelerator produces all energies up to 8 GeV
- Shielding is needed
- Beam-correlated neutrons mimic neutrino signal

#### Neutron Scatter on <sup>40</sup>Ar

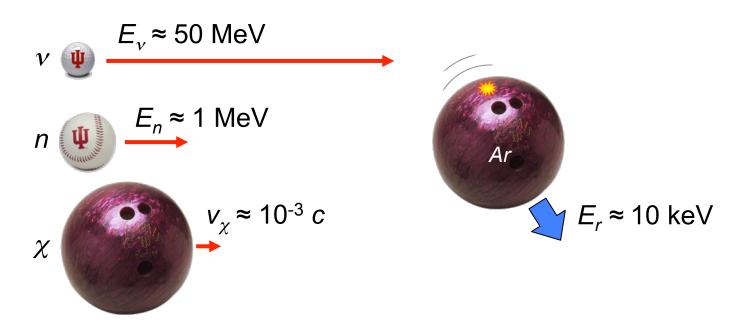


$$E_r^{\text{max}} = \frac{4\mathcal{M}}{(\mathcal{M}+1)^2} E_n \simeq 0.1 E_n$$

where 
$$\mathcal{M} = M/m_n$$

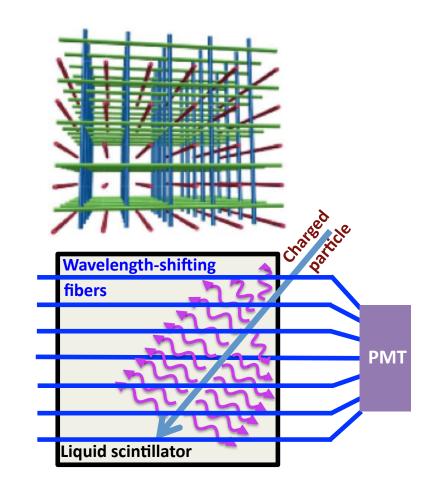
## Elastic Scattering Connection: ν, n, χ

- All these particle cause elastic scattering on argon
- Indistinguishable signal → ~10 keV nuclear recoil



#### **SciBath Detector**

- 80 L open volume of mineral oil based liquid scintillator
- Neutrons recoil off protons, create scintillation
- 768 wavelength shifting fibers readout
- IU built custom digitizer: 12 bit, 20 MS / s



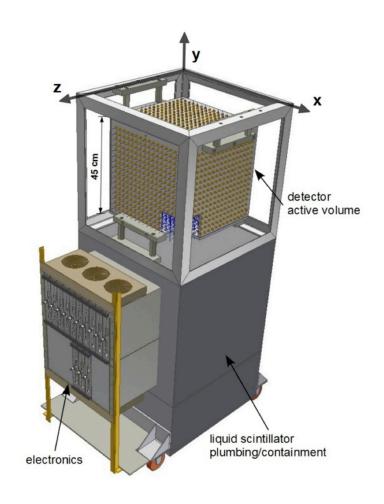
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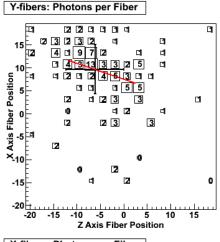


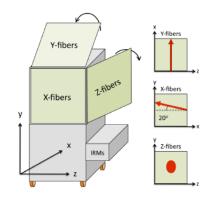
#### SciBath Detector

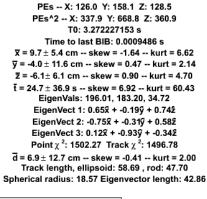
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### Sample Muon Candidate Event

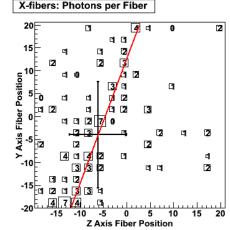


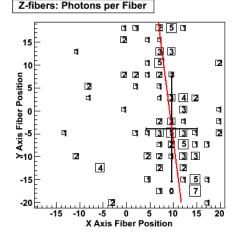


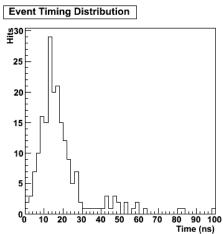


Event Num: 109 (1206) Multiplicity: 204

Total PEs: 412.6

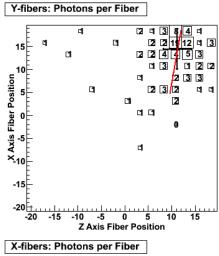


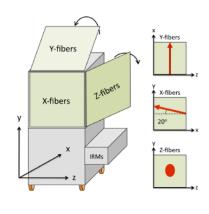




FNAL Neutrino Seminar -- R.L. Cooper

### Sample Neutron Candidate Event





3

3

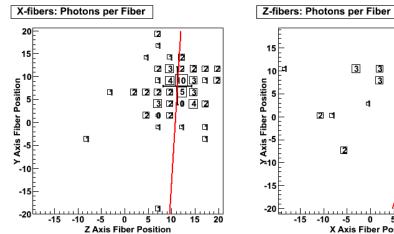
X Axis Fiber Position

[2] c1/

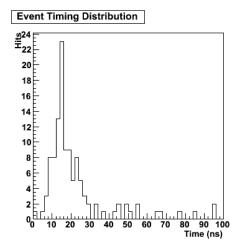
0 0 0 0 2 0 0

뎨

5 10 15 20



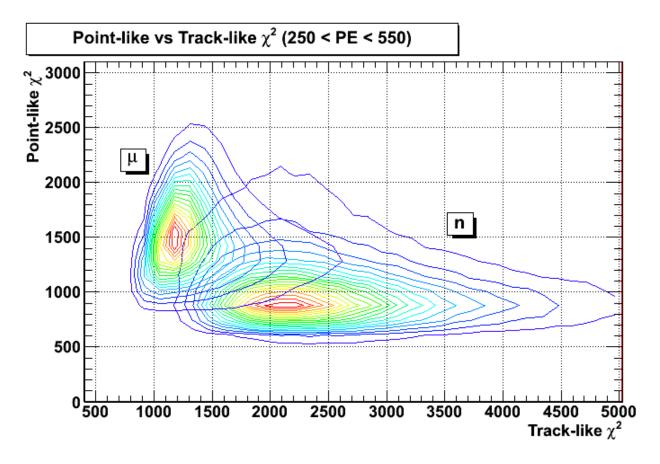




FNAL Neutrino Seminar -- R.L. Cooper

[2] [1

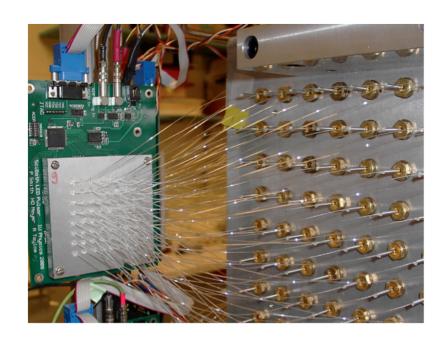
## $n / \mu$ Particle Discrimination



### Calibrating the SciBath Detector



- Low-light LED pulser (Y→Z)
- Use cosmic rays with known energy deposit (X→Y) requires previous calibration to count photons
- Detect 6 PEs / MeV→ want to improve

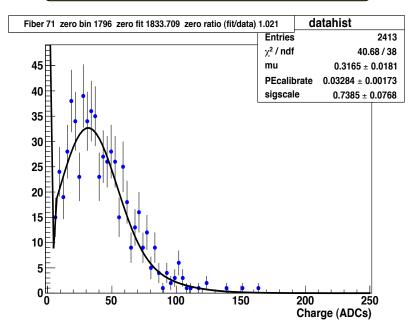


### Calibrating the SciBath Detector



- Low-light LED pulser (Y→Z)
- Use cosmic rays with known energy deposit (X→Y) requires previous calibration to count photons
- Detect 6 PEs / MeV
   → want to improve

#### **Single-PE LED Calibration**

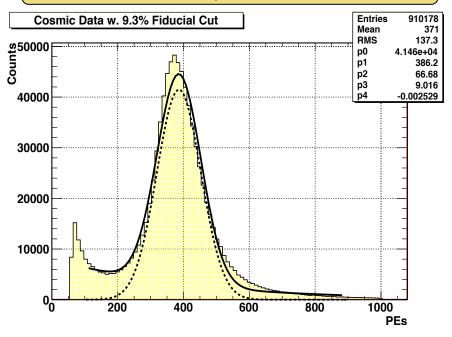


### Calibrating the SciBath Detector



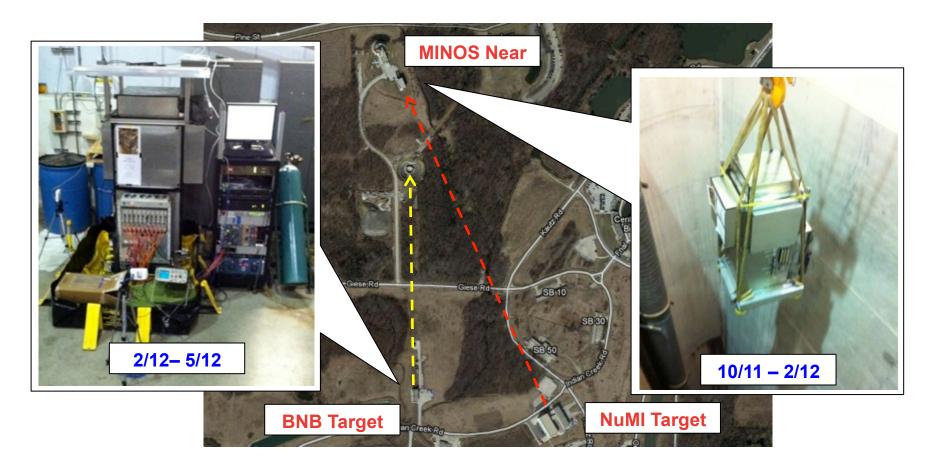
- Low-light LED pulser (Y→Z)
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#### **MIP Cosmic Ray Calibration Peak**



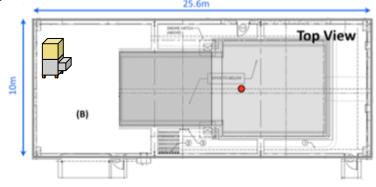


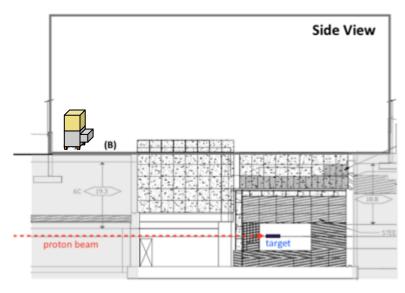
#### **Fermilab Measurement Sites**



### MI-12 Neutron Background Run

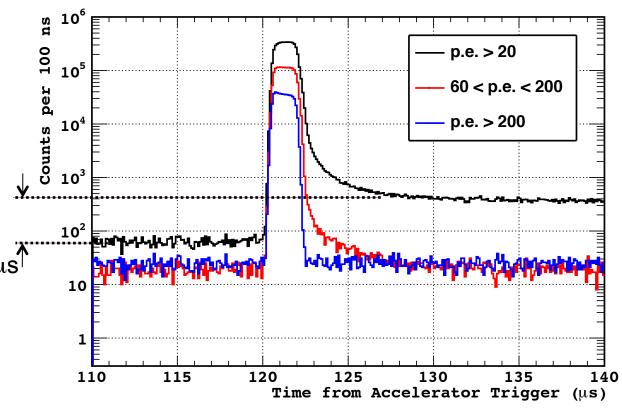
- Neutron flux ~20 m from target
- In-line behind beam target (ground)
- 29 Feb. 23 Apr., 2012
- 4.9x10<sup>19</sup> total protons on target (POT) (4.5x10<sup>12</sup> per pulse)





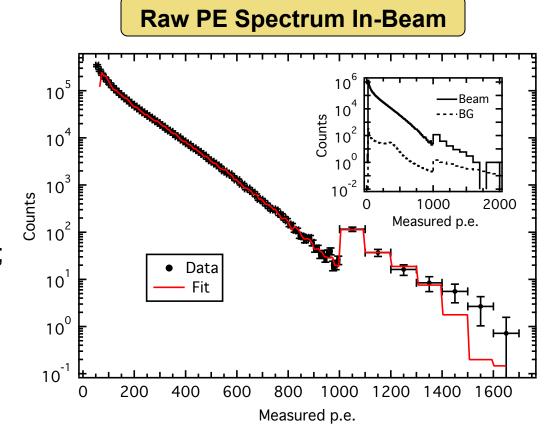
### MI-12 Beam Time Per PE "Group"

- HIGH PE group shows beam time structure
- MEDIUM PE group has few-µs excess – slower neutrons arriving later
- LOWEST PE group has significant excess 200 μs lifetime from n(p, d)γ neutron capture reaction



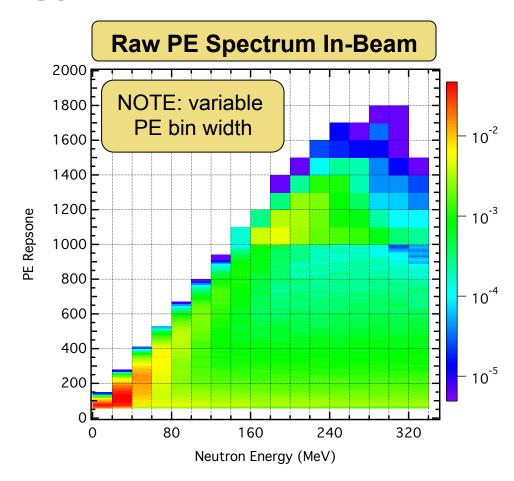
### **BNB Neutron Energy Spectrum**

- E<sub>n</sub> unfolded from PEs spectrum simulation of detector response
- 2.44 ± 0.34 pulse<sup>-1</sup> m<sup>-2</sup>  $(E_n > 40 \text{ MeV})$
- Lose sensitivity > 200 MeV;
- Neutron spectrum
   20 m from BNB



### **BNB Neutron Energy Spectrum**

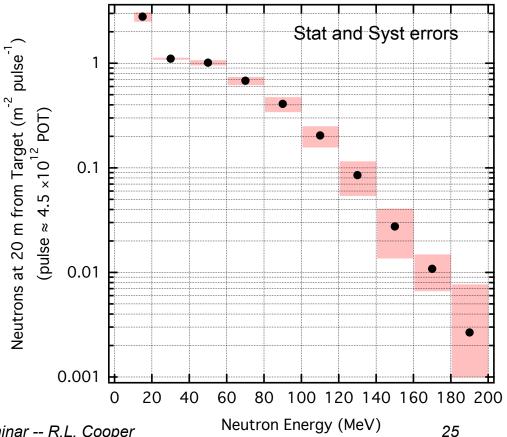
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### **BNB Neutron Energy Spectrum**

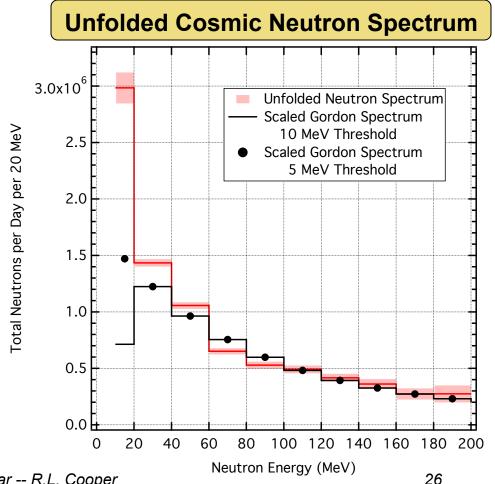
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- $2.44 \pm 0.34 \text{ pulse}^{-1} \text{ m}^{-2}$  $(E_n > 40 \text{ MeV})$
- Lose sensitivity > 200 MeV;
- Neutron spectrum 20 m from BNB

#### **Unfolded Neutron Energy Spectrum**



## Validation of Unfolding Techniques

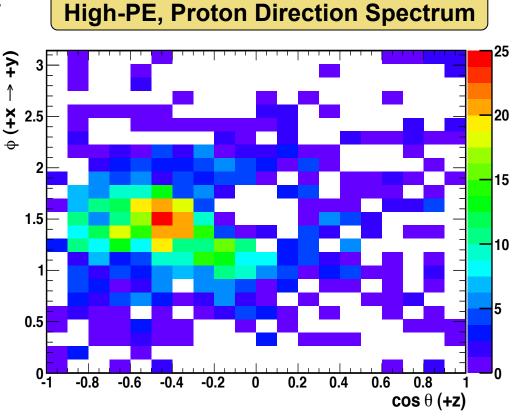
- Cosmic ray neutron spectrum also unfolded
- Gordon et al., IEEE TNS 51, (2004) 3427 parameterizes surface neutron flux from Bonner sphere data
- Energy shape matches, overall scale factor needed



### **Direction Spectrum**

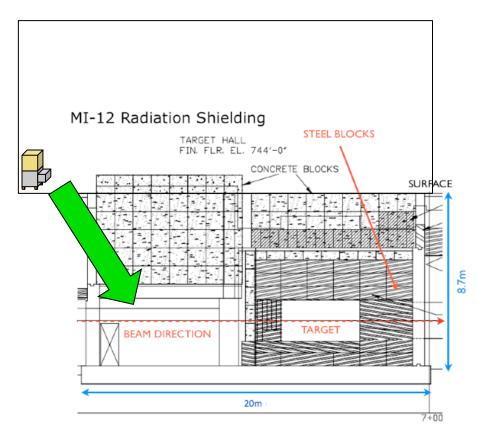
 High PE protons will be tracklike; can be imaged

- Principle component analysis yields eigenvector
- Back-projecting direction spectrum tends to point upstream of target
- Tracking validated with cosmic rays and NuMI beam



### **Direction Spectrum**

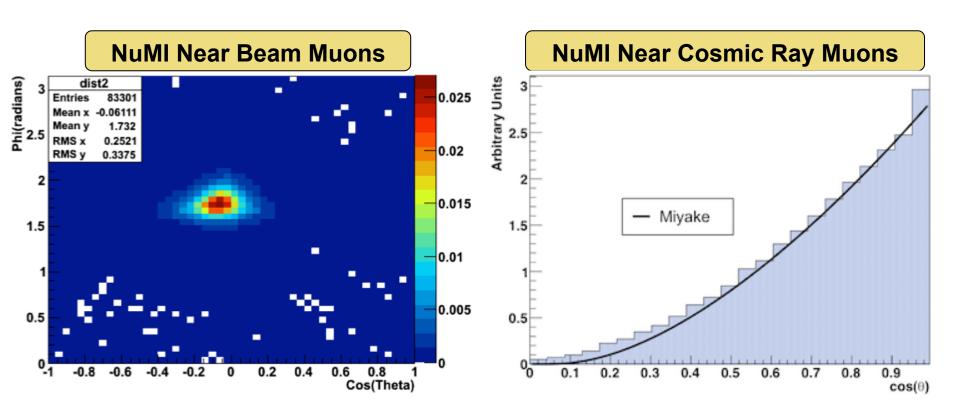
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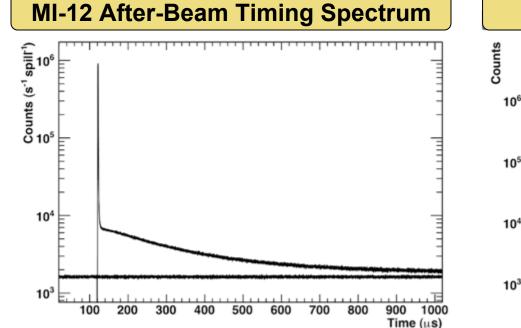
## Validation of SciBath Tracking

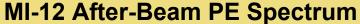
Tracking algorithms validated with NuMI underground data

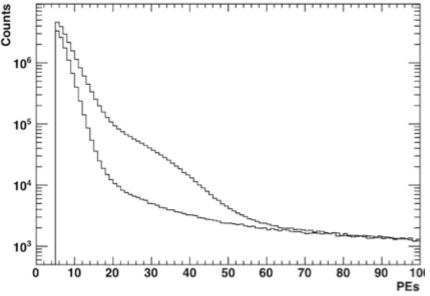


## **Capture-Gated Neutrons at MI-12**

• At surface, accidental rate (and high primary rate) precludes  $n(p,d)\gamma$  capture gating  $\rightarrow$  clear statistical sensitivity to thermal captures



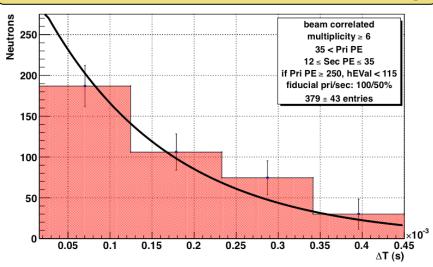




## Capture-Gated Neutrons at NuMI

• NuMI near hall (100 m overburden) capture-gating neutron spectroscopy technique demonstrated (L. Garrison thesis)

#### **NuMI Beam Neutron Capture Timing**

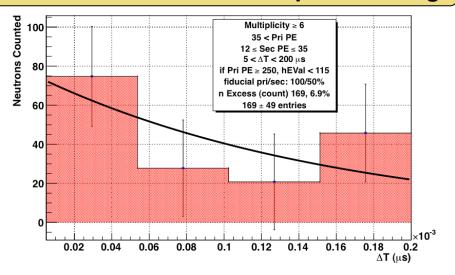


#### **NuMI Beam Neutron Energy** Primary in beam ×10<sup>-6</sup> Tagged n Capture Sec Neutron Flux (cm<sup>-2</sup> s<sup>-1</sup> MeV<sup>-1</sup>) 0. 1. 5. Multiplicity ≥ 6 Pri PE > 35 12.5 < Sec PE < 35 If Pri PE>250, hEVal<115 16 < ΔT < 450 μs Fiducial pri/sec: 100/50% **BG Subtracted** 366 ± 41 Events 0.1 0.05 20 40 60 80 100 120 MeV

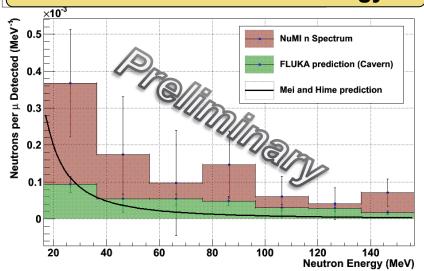
## Capture-Gated Neutrons at NuMI

• NuMI near hall (100 m overburden) capture-gating neutron spectroscopy technique demonstrated (L. Garrison thesis)

#### **NuMI Cosmic Neutron Capture Timing**

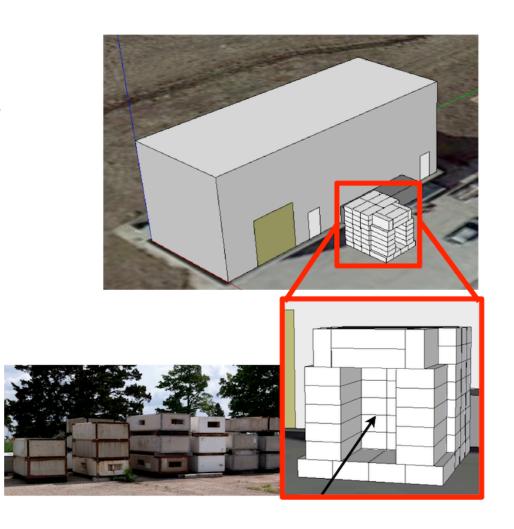


#### **NuMI Cosmic Neutron Energy**



#### **Current Studies**

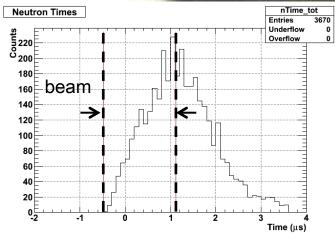
- 2012 measurements at one position with no shielding
- We are improving SciBath, building concrete shielding
- Locate a viable location for CENNS & CAPTAIN
- Survey the area with portable detector



#### **Current Studies**

- 2012 measurements at one position with no shielding
- We are improving SciBath, building concrete shielding
- Locate a viable location for CENNS & CAPTAIN
- Survey the area with portable detector







## **Beam Off-Target Rates (> 0.5 MeV)**

#### 50 m Absorber

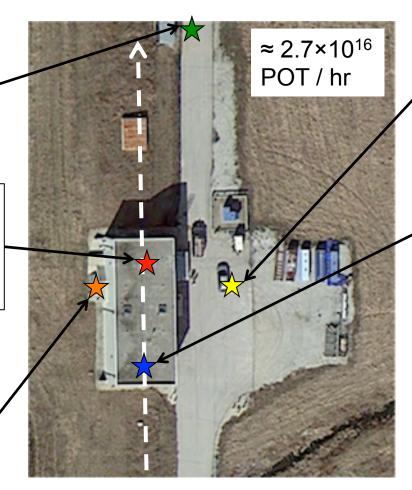
- 6 m from Fe beam stop
- 310 n / 10<sup>16</sup> POT

#### **Collimator**

- 8 m from Be beam target
- 5608 n / 10<sup>16</sup> POT

#### **Stairwell**

- 9 m from Be beam target
- 1384 n / 10<sup>16</sup> POT



#### Target 90° FOX

- 20 m from Be beam target
- 390 n / 10<sup>16</sup> POT

#### 2012 SciBath Loc

- 20 m from Be beam target
- 211 n / 10<sup>16</sup> POT

Neutron spectrum unfolding underway

#### **CENNS-10**

#### Goals

- Develop LAr technology
- Perform very high-energy neutron calibrations

#### **Status**

- Moved to Indiana
- Planning for calibration at Los Alamos WNR neutron beam



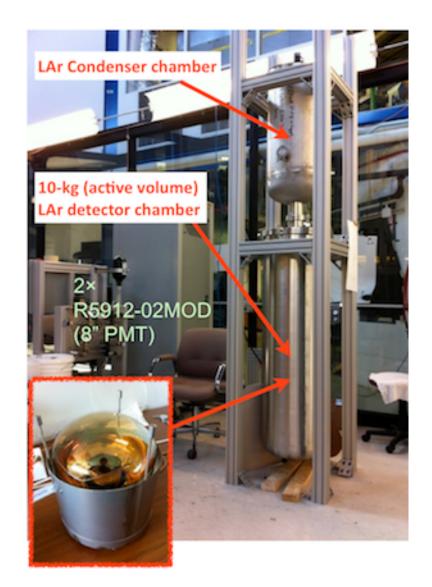
#### **CENNS-10**

#### Goals

- Develop LAr technology
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#### **Status**

- Moved to Indiana
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### **Summary of BNB Work for CENNS**

#### **SciBath**

Fast neutron measurements (10-200 MeV)





**MiniCLEAN** 

First CENNS measurement





#### EJ-301 Cells

Portable array (0.5-20 MeV)



preparatory measurements

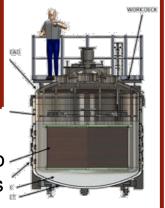


#### CENNS-10

10 kg LAr testing prototype

# **CAPTAIN**

Low-E neutrino cross sections



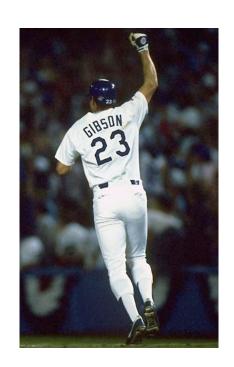
#### **Summer 2015 Plans**

- BNB: Plan to measure near BNB target building for CENNS, CAPTAIN, and general SBN program (May or June for 1 month)
- SciBooNE: Measure high-energy neutrino-induced neutrons and constrain thermal neutron rates from n(p,d)γ capture rates:

relevant for ANNIE, microBooNE, and SBN (May *or* June for 1 month)



## **BACKUPS**



## Structure of the CENNS Signal

Predicted scattering rate

$$\frac{d\sigma}{dE} = \frac{G_F^2}{4\pi} \left[ \frac{(1 - 4\sin^2\theta_w)Z - N}{2} M \left( 1 - \frac{ME}{2E_\nu^2} \right) F(Q^2)^2 \right]$$

Detection Rate [ton-1 year-1]

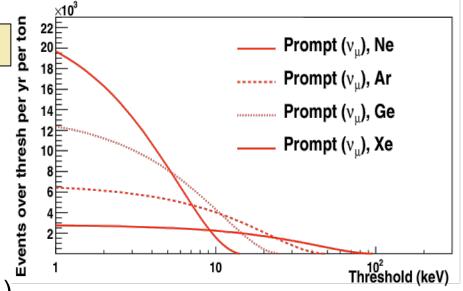
≈ 0 → protons have little influence

square of sum → part of coherence condition

nuclear form factor

→ distribution of neutrons

Recoil energy  $(M^{-1})$  and rate  $(N^2)$ 



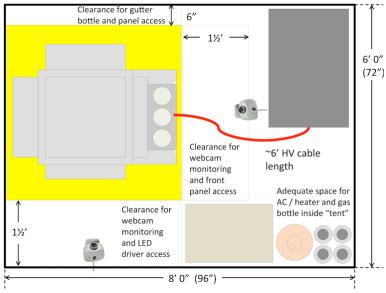
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- 29 Feb. 23 Apr., 2012
- 4.9x10<sup>19</sup> total protons on target (POT) (4.5x10<sup>12</sup> per pulse)



### **Utility Trailer for BNB Measurement**







### **CENNS-10 On the Move**







## **Off-Target Runs**

